

LECTURES ABOUT (ADVANCED) STATISTICAL PHYSICS

T.S.Biró, MTA Wigner Research Centre for Physics, Budapest

Lectures given at: University of Johannesburg, South-Africa,

November 26 – November 29, 2012.

- 1. Ancient Thermodynamics (... - 1870)**
- 2. The Rise of Statistical Physics (1890 – 1920)**
- 3. Modern (postwar) Problems (1940 – 1980)**
- 4. Corrections (1950 – 2005)**
- 5. Generalizations (1960 – 2010)**
- 6. High Energy Physics (1950 – 2010)**

LECTURE FOUR ABOUT (ADVANCED) STATISTICAL PHYSICS

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Generalized kinetic theory



Simulation using non-additive rule

PRL 95: 162302, 2005

with Gábor Purcsei

Non-extensive Boltzmann Equation

(NEBE) with Gábor Purcsei:

- Rényi-Tsallis energy addition rule
- random momenta accordingly
- pairwise collisions repeated
- momentum distribution collected

Boltzmann algorithm: pairwise combination + separation

With additive composition rule at independence:

$$f_1 \cdot f_2 = f_3 \cdot f_4$$

$$E_1 + E_2 = E_3 + E_4$$

$$f_{\text{eq}} = \frac{1}{Z} e^{-\beta E}$$

Such rules generate exponential distribution

Boltzmann algorithm: pairwise combination + separation

With associative composition rule at independence:

$$\mathbf{f}_1 \cdot \mathbf{f}_2 = \mathbf{f}_3 \cdot \mathbf{f}_4$$

$$E_1 \oplus E_2 = E_3 \oplus E_4$$

$$L(E_1) + L(E_2) = L(E_3) + L(E_4)$$

$$\mathbf{f}_{\text{eq}} = \frac{1}{Z} e^{-\beta L(E)}$$

Such rules generate 'exponential of the formal logarithm' distribution

Generalized Stoßzahlansatz

$$DF(f_1) = \int_{234} w_{1234} (G_{34} - G_{12})$$

$$DF = \frac{p^\mu}{p^0} \partial_\mu F$$

$$G_{ij} = e_a \left(\ln_a(f_i) + \ln_a(f_j) \right)$$

General H theorem

$$S^\mu = - \int \frac{p^\mu}{p^0} \sigma(F(f))$$

$$\partial_\mu S^\mu = - \int_1 \sigma'(F) DF(f_1) =$$

$$\frac{1}{4} \int_{1234} w_{1234} (\sigma'_1 + \sigma'_2 - \sigma'_3 - \sigma'_4) (G_{12} - G_{34})$$

$$\partial_\mu S^\mu \geq 0 \quad \text{iff} \quad \Phi(G_{ij}) = \sigma'_i + \sigma'_j$$

a monotonic rising function

General H theorem: entropy density formula

$$\Phi(G) = \alpha \ln_a(G) + 2\beta$$

$$\sigma'(F(f)) = \alpha \ln_a(f) + \beta$$

$$\sigma(f) = \int F'(f)(\alpha \ln_a(f) + \beta)df$$

Detailed balance: $G_{12} = G_{34}$

Stoßzahlen satz :

$$G_{ij} = e_a \left(\ln_a f_i + \ln_a f_j \right) = e^{L_a \left(L_a^{-1} (\ln f_i) + L_a^{-1} (\ln f_j) \right)}$$

Conserved in microcollision :

$$X(E_1) + X(E_2) = X(E_3) + X(E_4)$$

Detailed balance: $G_{12} = G_{34}$

General solution :

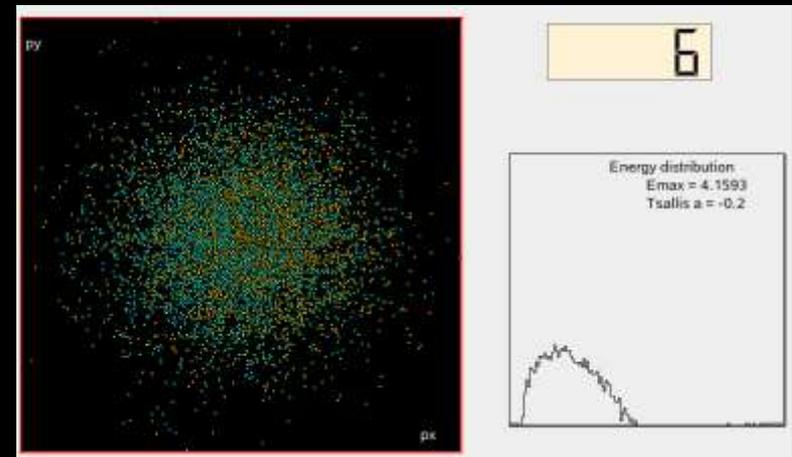
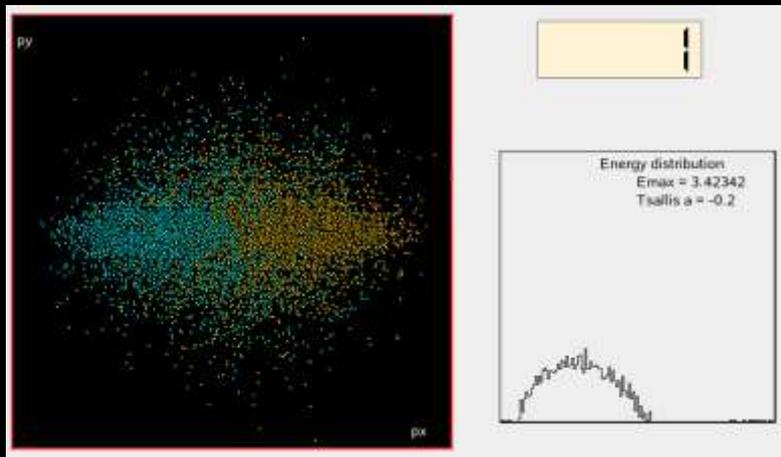
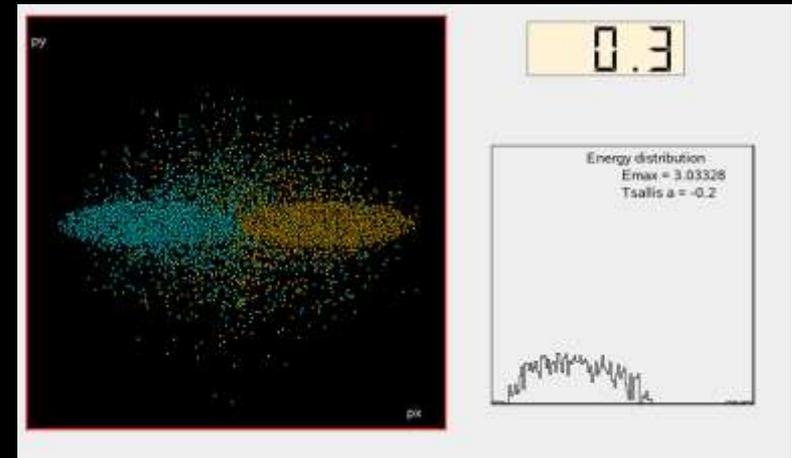
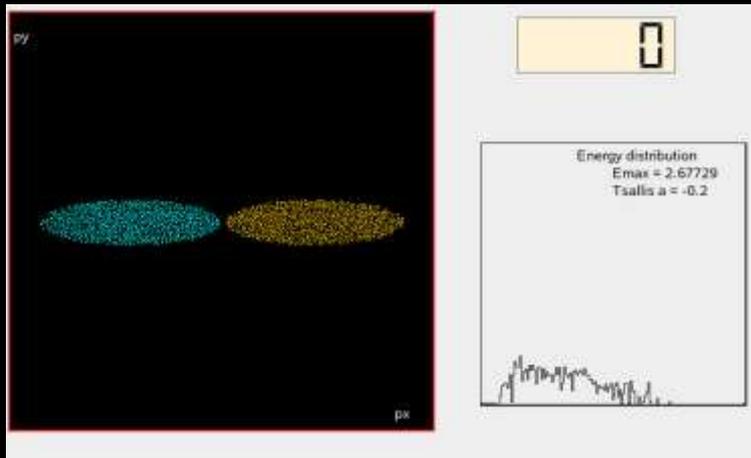
$$L_a^{-1}(\ln f_i) = \alpha - \beta X(E_i)$$

$$f_i = \exp(L_a[\alpha - \beta X(E_i)])$$

Detailed balance: proof

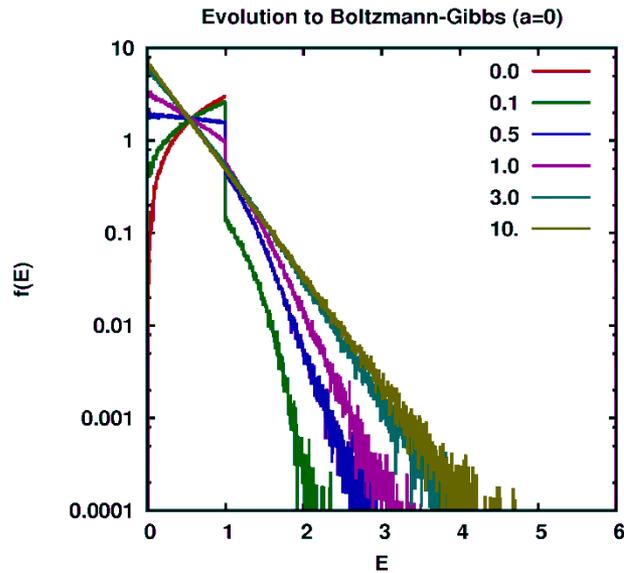
$$\begin{aligned} G_{12} &= \exp\left(L_a \left[L_a^{-1}(\ln f_1) + L_a^{-1}(\ln f_2) \right]\right) \\ &= \exp\left(L_a \left[2\alpha - \beta X(E_1) - \beta X(E_2) \right]\right) \\ &= \exp\left(L_a \left[2\alpha - \beta X(E_3) - \beta X(E_4) \right]\right) \\ &\exp\left(L_a \left[L_a^{-1}(\ln f_3) + L_a^{-1}(\ln f_4) \right]\right) = G_{34} \end{aligned}$$

Evolution in NEBE phase space

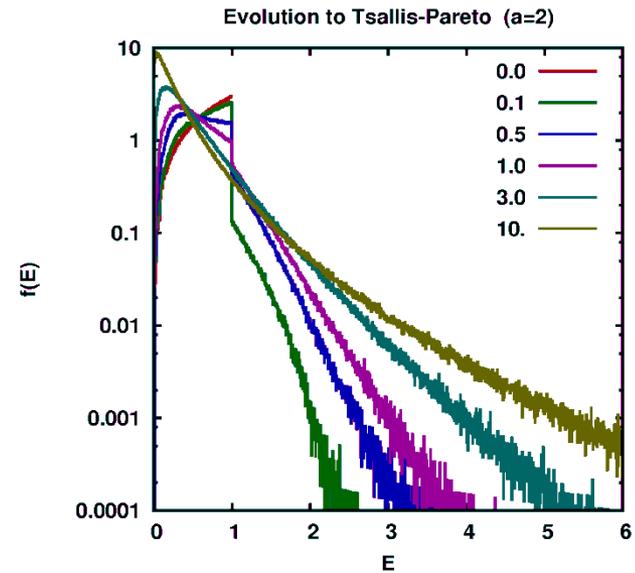


Stationary energy distributions in NEBE program

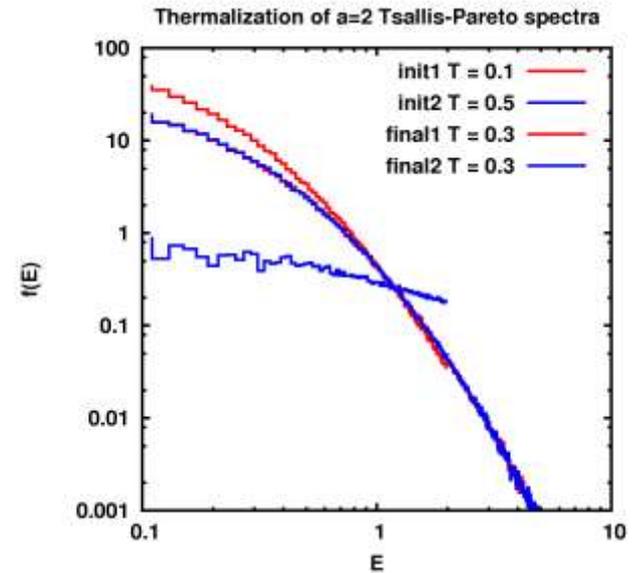
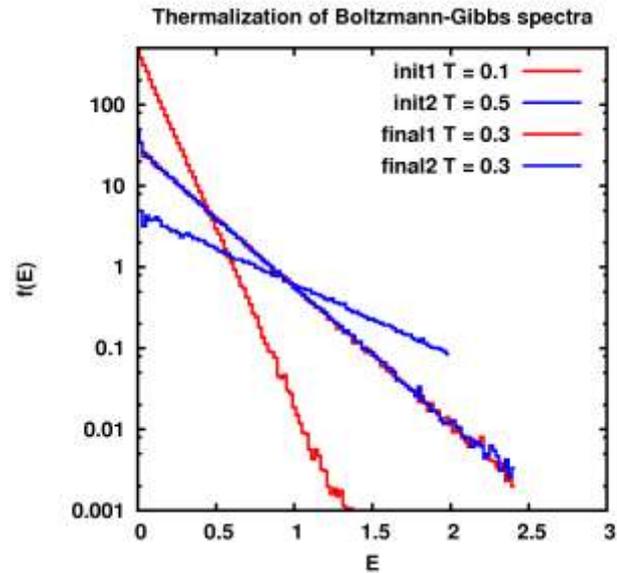
$x + y$



$x + y + 2xy$



Thermal equilibration in NEBE program



Microscopic theory in non-extensive approach: questions, projects, ...

- **Ideal gas with deformed exponentials**
- **Boltzmann and Bose distribution**
- **Fermi distribution: pti – hole effect**
- **Thermal field theory with stochastic temperature (Euler-Gamma distribution)**
- **Lattice SU(2) with non-ext. Metropolis method**

Ideal Tsallis-Bose gas

$$\ln Z = V \dots \int E^2 dE \left(-\ln(1 - e_{-a}(-\beta E)) \right)$$

$$-\frac{1}{V} \frac{\partial \ln Z}{\partial \beta} = \dots \int E^2 dE \frac{E}{1 + a\beta E} \frac{1}{e_a(\beta E) - 1}$$

$$e = \frac{\gamma T^4}{2\pi^2} \sum_{n=1}^{\infty} \frac{3!}{n(n-a)(n-2a)(n-3a)}$$

For $c = 5.5$ we have $1/a = 4.5$ and $e \approx 4 e_0$

Limiting temperature with Tsallis distribution

PLB 632:247,2006

(with A. Peshier, Giessen)

Massless particles, 3-dim. momenta, N-fold

$$3T = \frac{E/N}{1 + a E/N}$$

Hagedorn

For $N \geq 4$: Tsallis partons \rightarrow Hagedorn states

If it were $a \sim 1/T \rightarrow$ there would not be a limiting T !!!

Tsallis lattice EOS

**Tamás S. Bíró (KFKI RMKI Budapest) and
Zsolt Schram (DTP ATOMKI Debrecen)**

- **Lattice action with superstatistics**
- **Ideal gas with power-law tails**
- **Numerical results on EOS**

Lattice theory

Expectation values of observables:

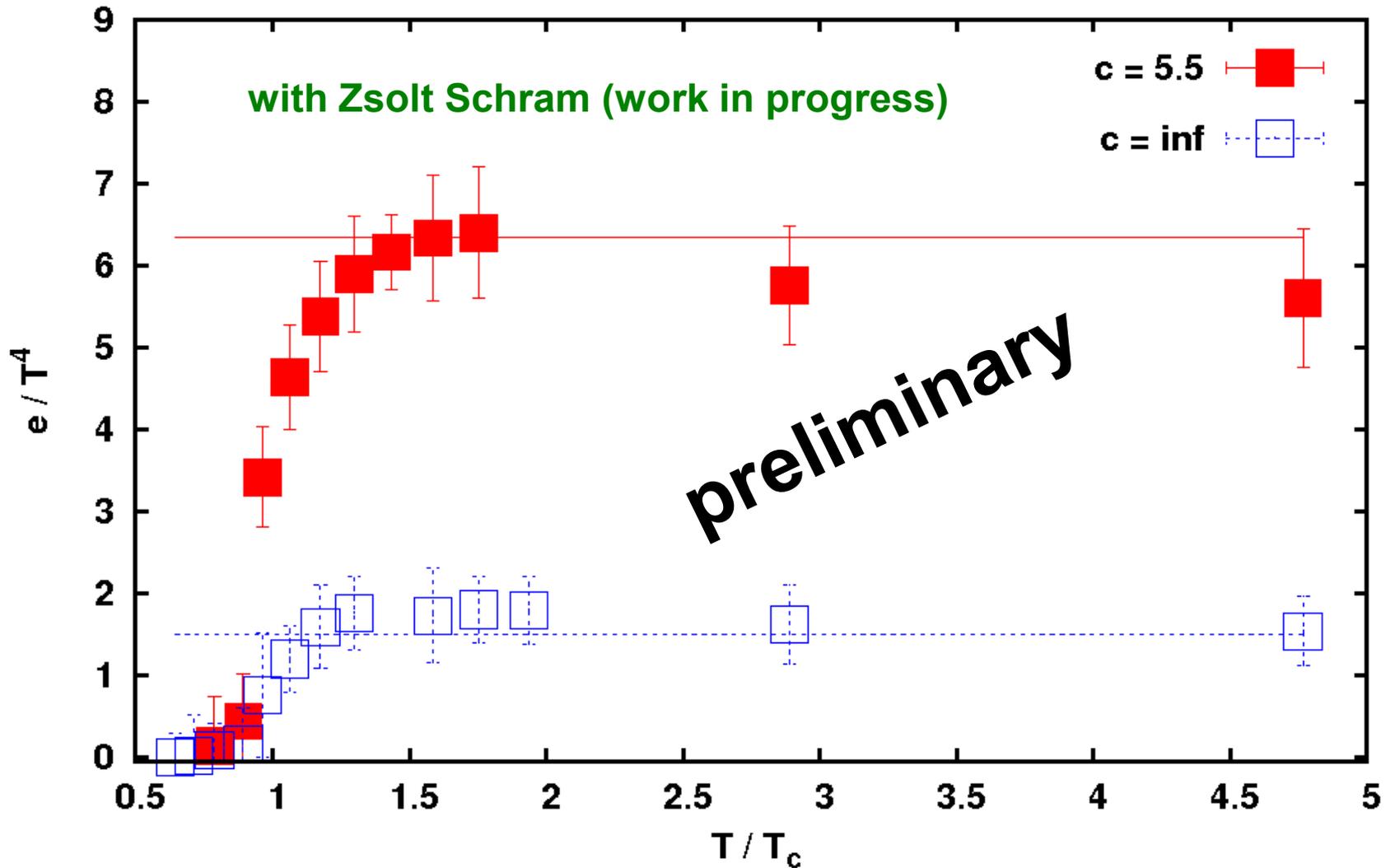
$$\langle A \rangle = \frac{\int DU \int dt w_{\zeta}(t) e^{-S(t,U)} A(U)}{\int DU \int dt w_{\zeta}(t) e^{-S(t,U)}}$$

Action: $S(t,U) = a(U) t + b(U) / t$

$t = a_t / a_s$ *asymmetry parameter*

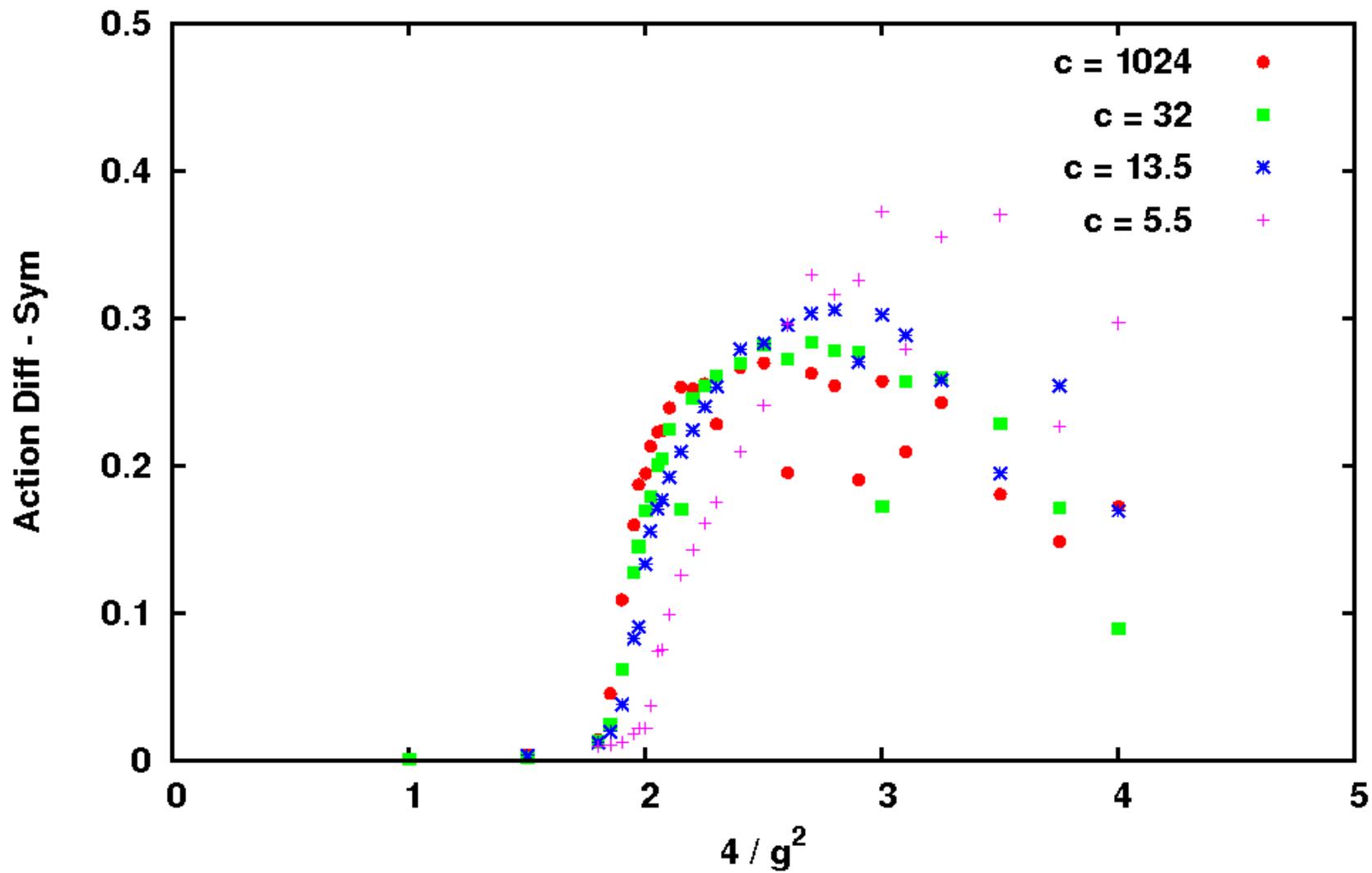
SU2 Yang-Mills eos on the lattice with Euler-Gamma distributed inverse temperature: Effective action method

SU2 Tsallis $c=5.5$ $c=\text{inf}$ energy density vs temperature



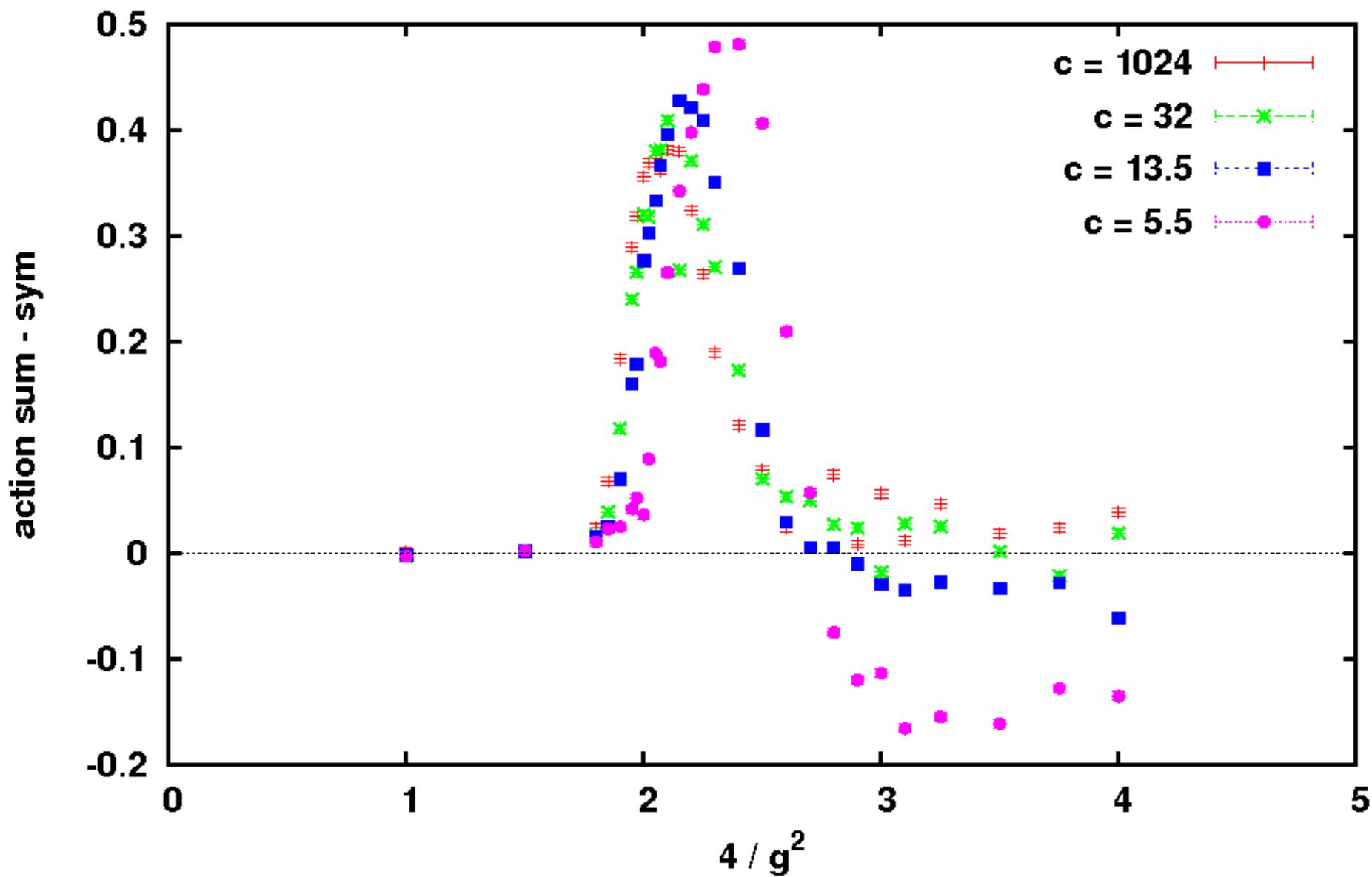
e / T4

TS su2 $10^4 - 10^3 \times 2$ lattice values



(e-3p) / T4

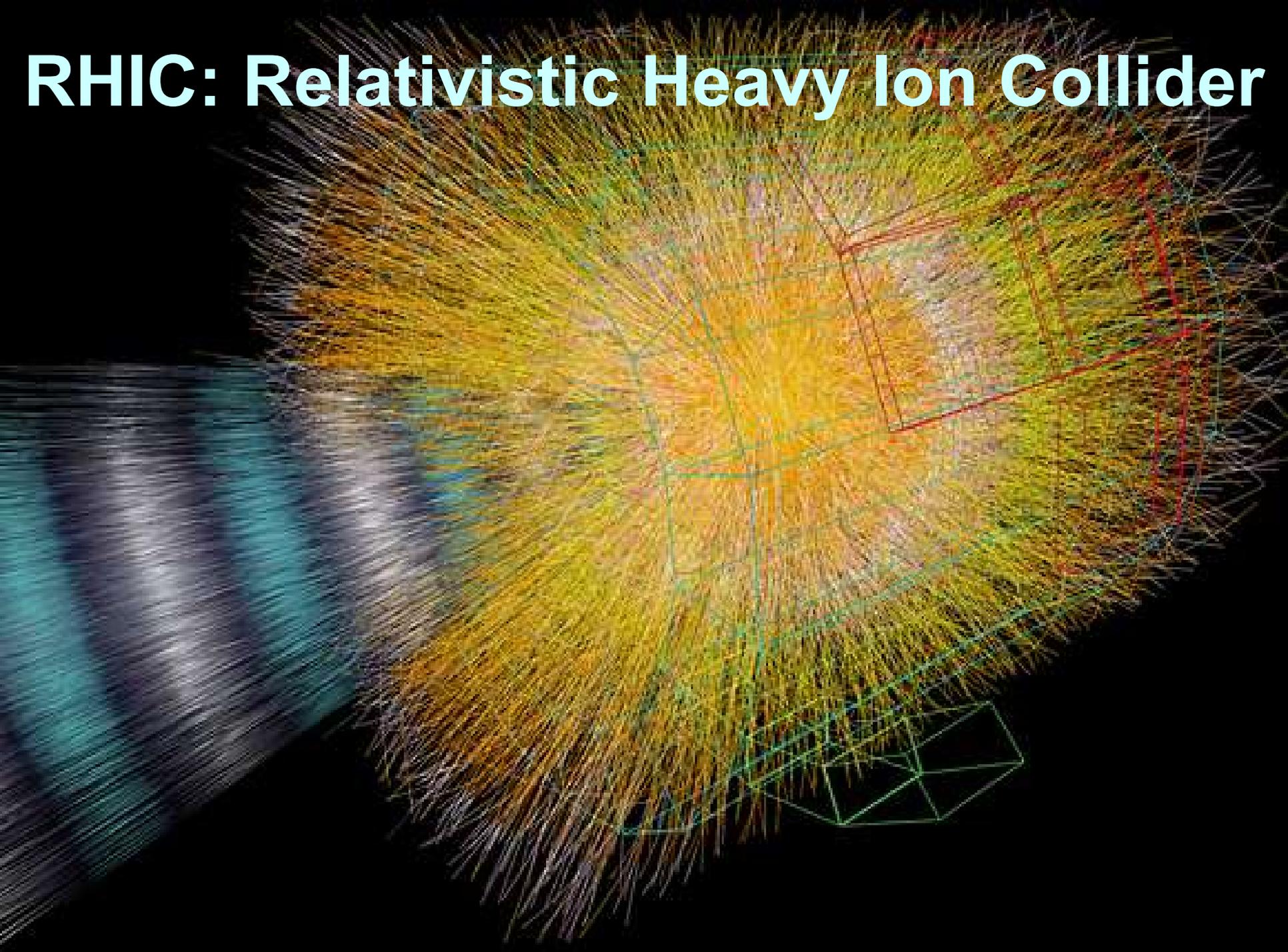
TS su2 $10^4 - 10^3 \times 2$ lattices



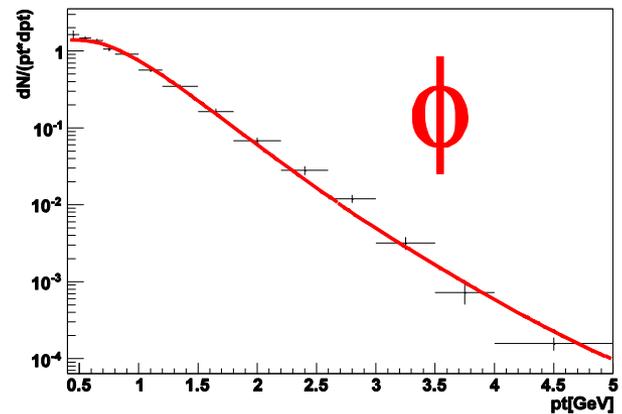
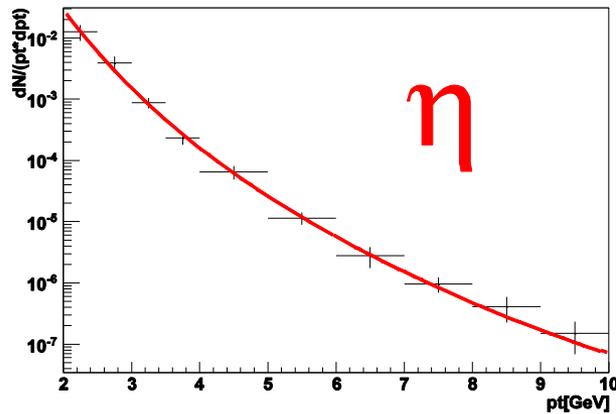
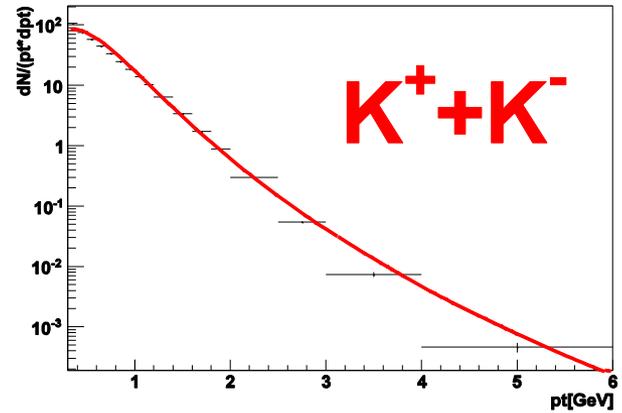
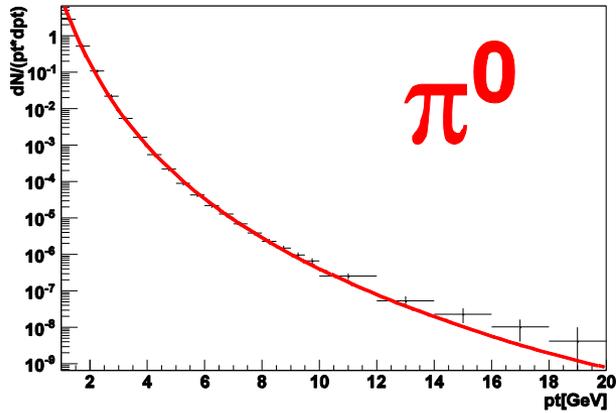
HIGH ENERGY PHYSICS

- **Old and New Statistical Models, Hydrodynamics**
- **Multiplicity Fluctuations**
- **Valence Quark Scaling**
- **Non-extensive Boltzmann Equation**
- **Fragmentation Statistics**
- **Classical Fields and Thermal Gravity**

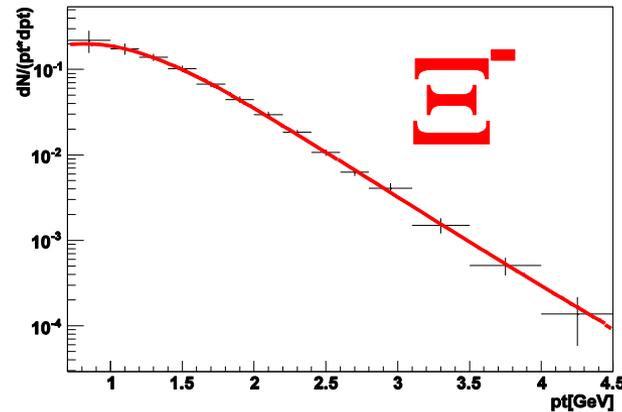
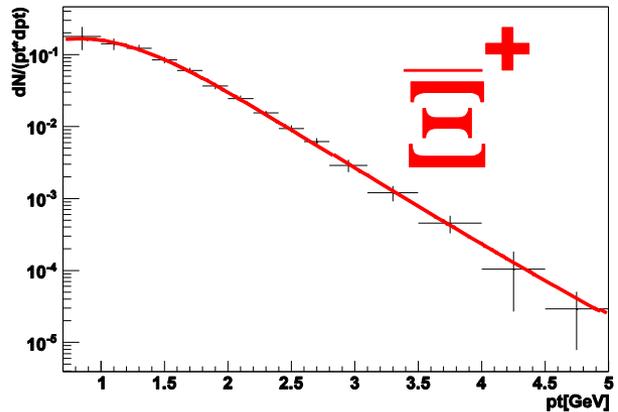
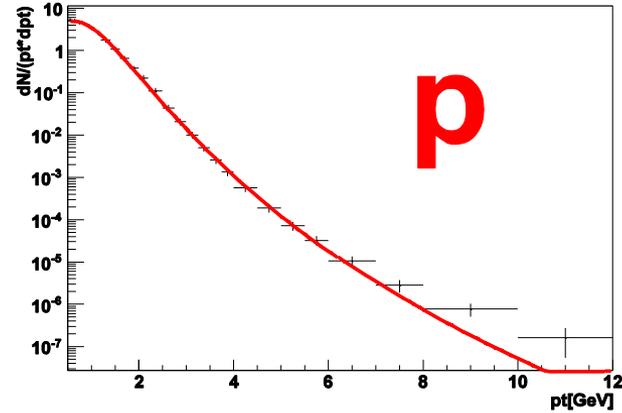
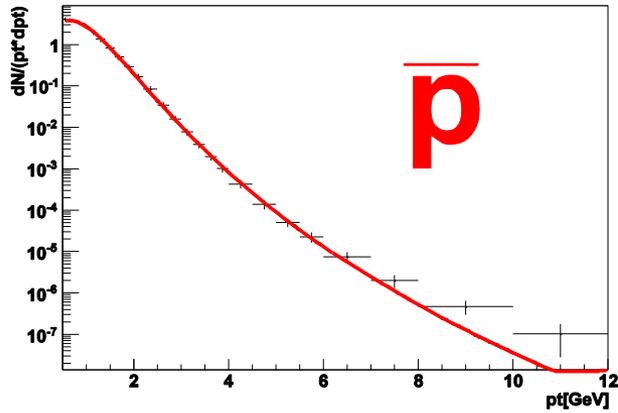
RHIC: Relativistic Heavy Ion Collider



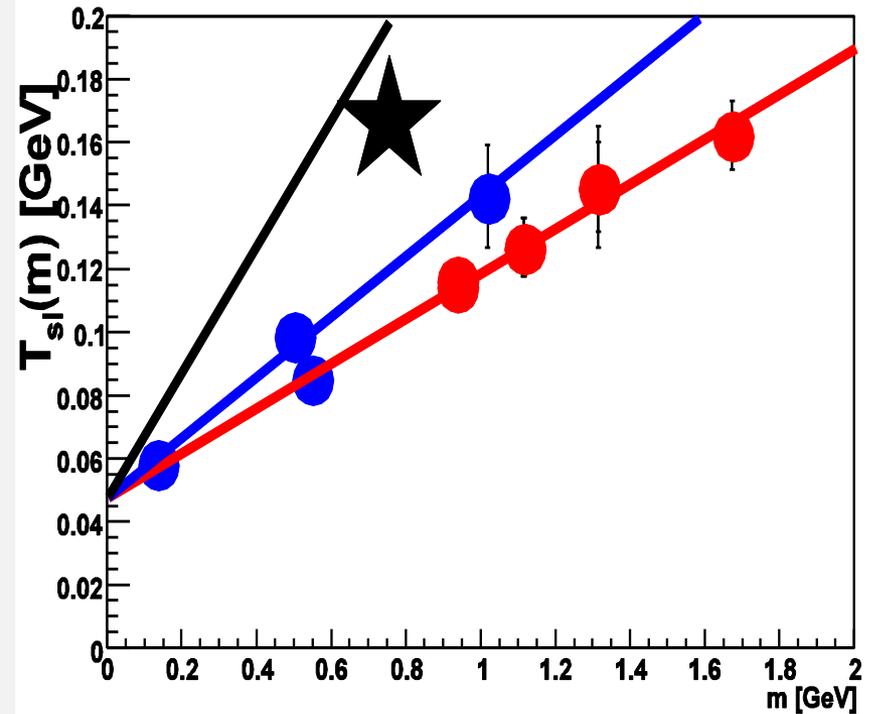
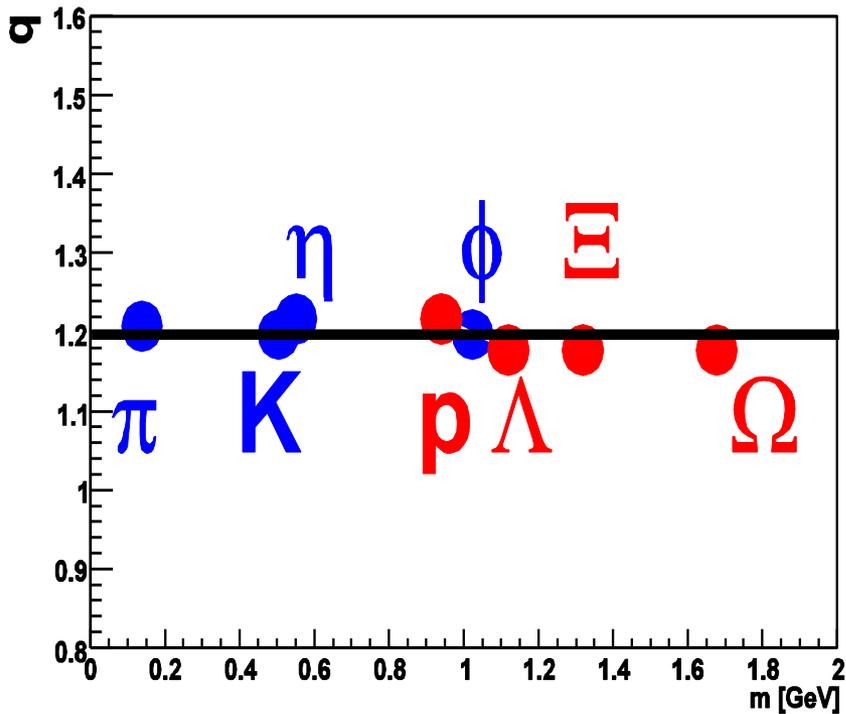
Tsallis quark matter + transverse flow + quark coalescence fits to hadron spectra



Tsallis quark matter + transverse flow + quark coalescence fits to hadron spectra

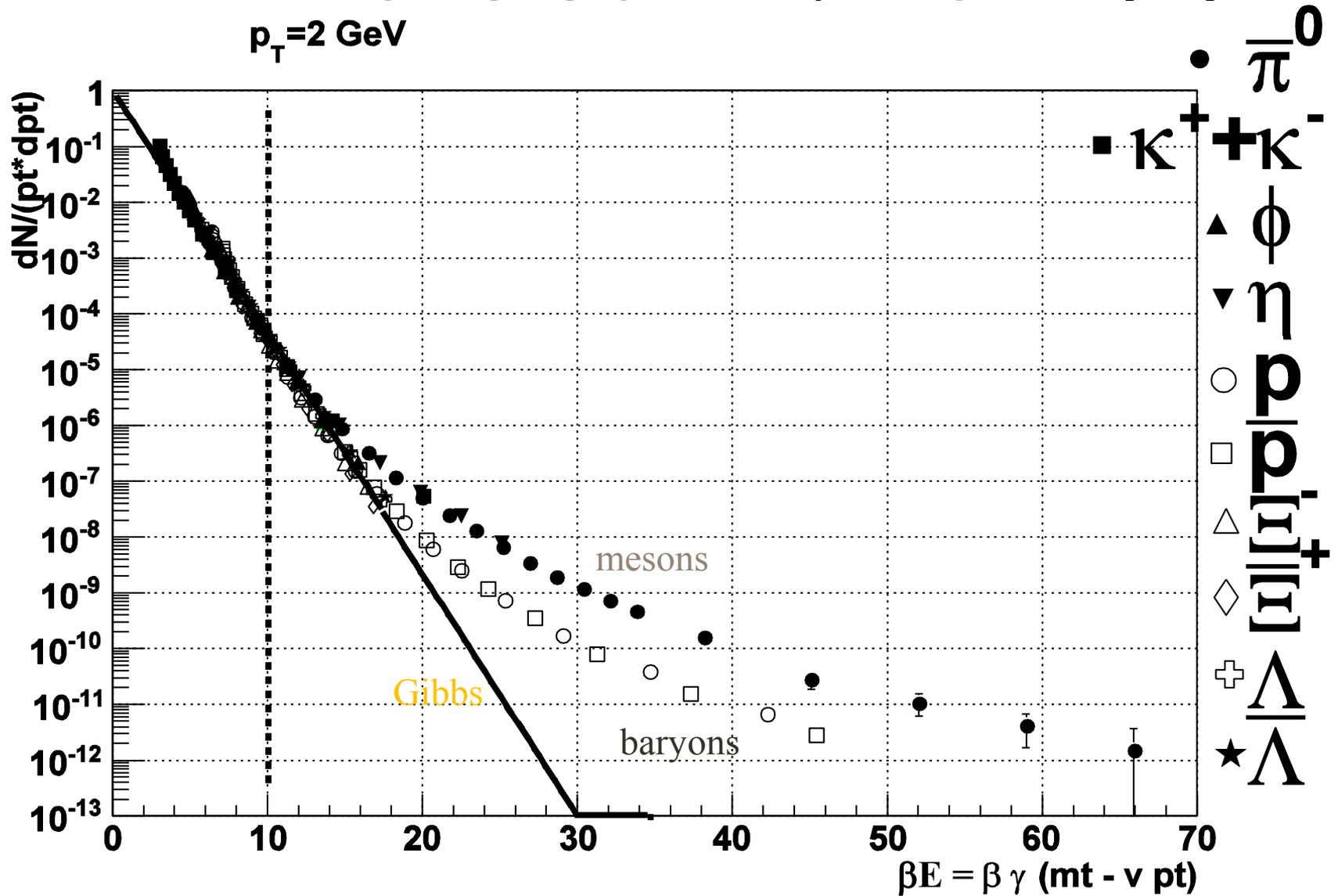


Blast wave fits and quark coalescence

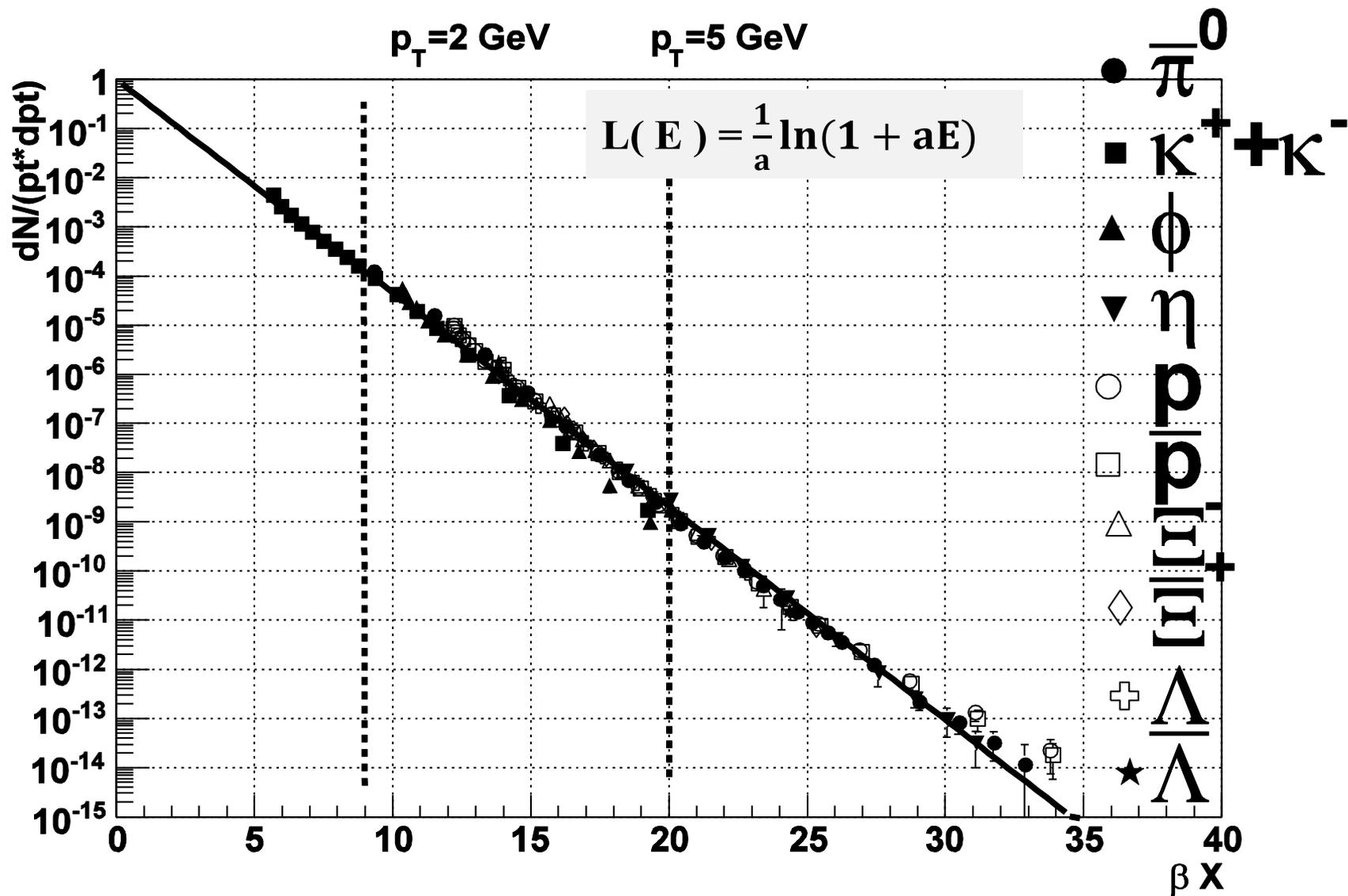


WHAT DOES SCALE, E OR L(E)?

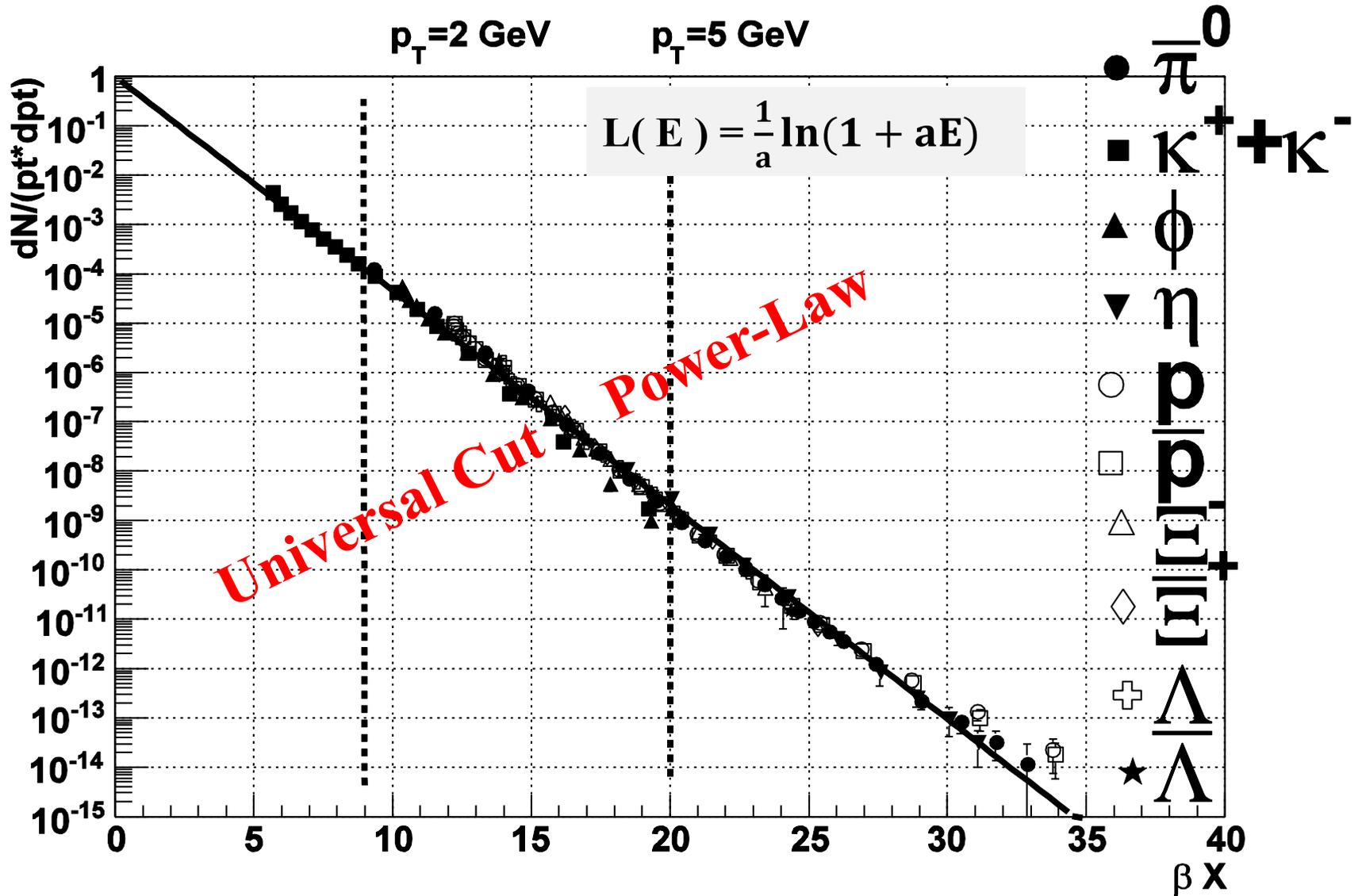
$p_T = 2$ GeV



WHAT DOES SCALE, E OR L(E)?



WHAT DOES SCALE, E OR L(E)?



LIMITING TEMPERATURE WITH TSALLIS DISTRIBUTION

(with A. Peshier, Giessen)

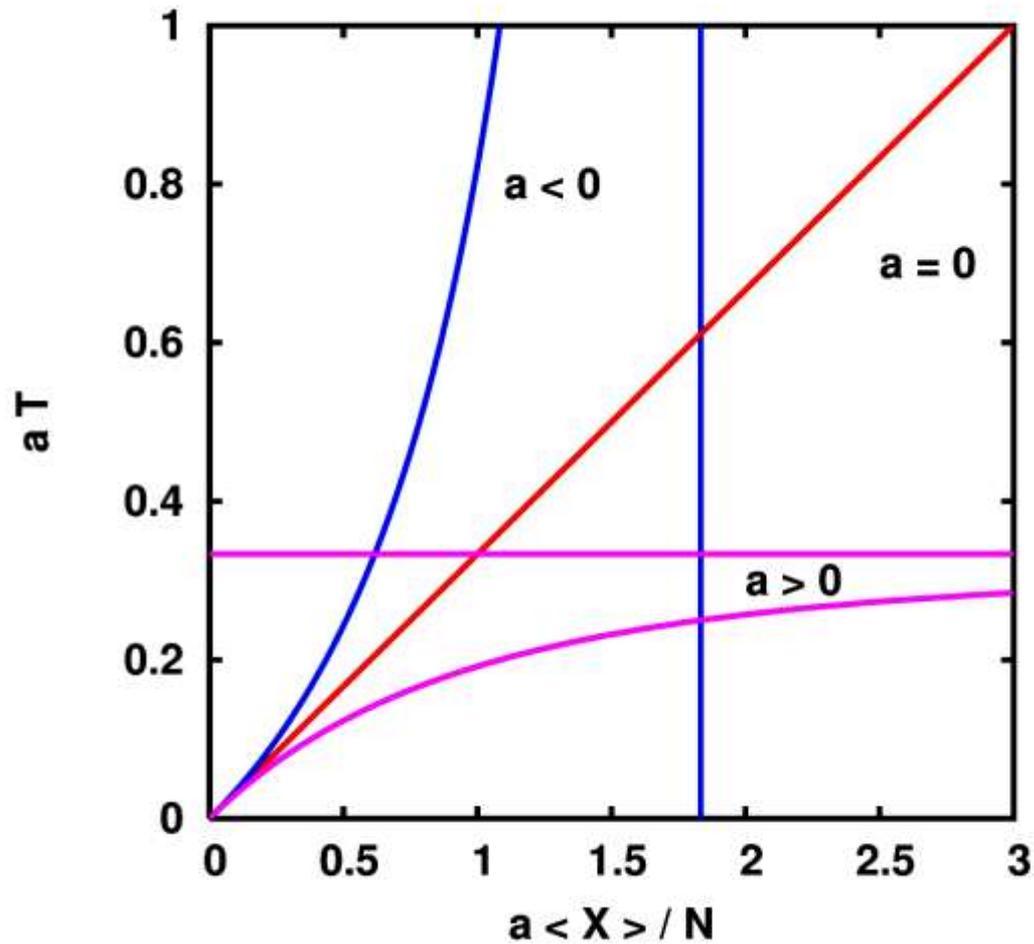
hep-ph/0506132, PLB

Massless particles, d-dim. momenta, N-fold

$$\frac{\langle X(E) \rangle}{N} = \sum_{j=1}^d \frac{T E_c}{E_c - j T} ; \quad T_H = E_c / d$$

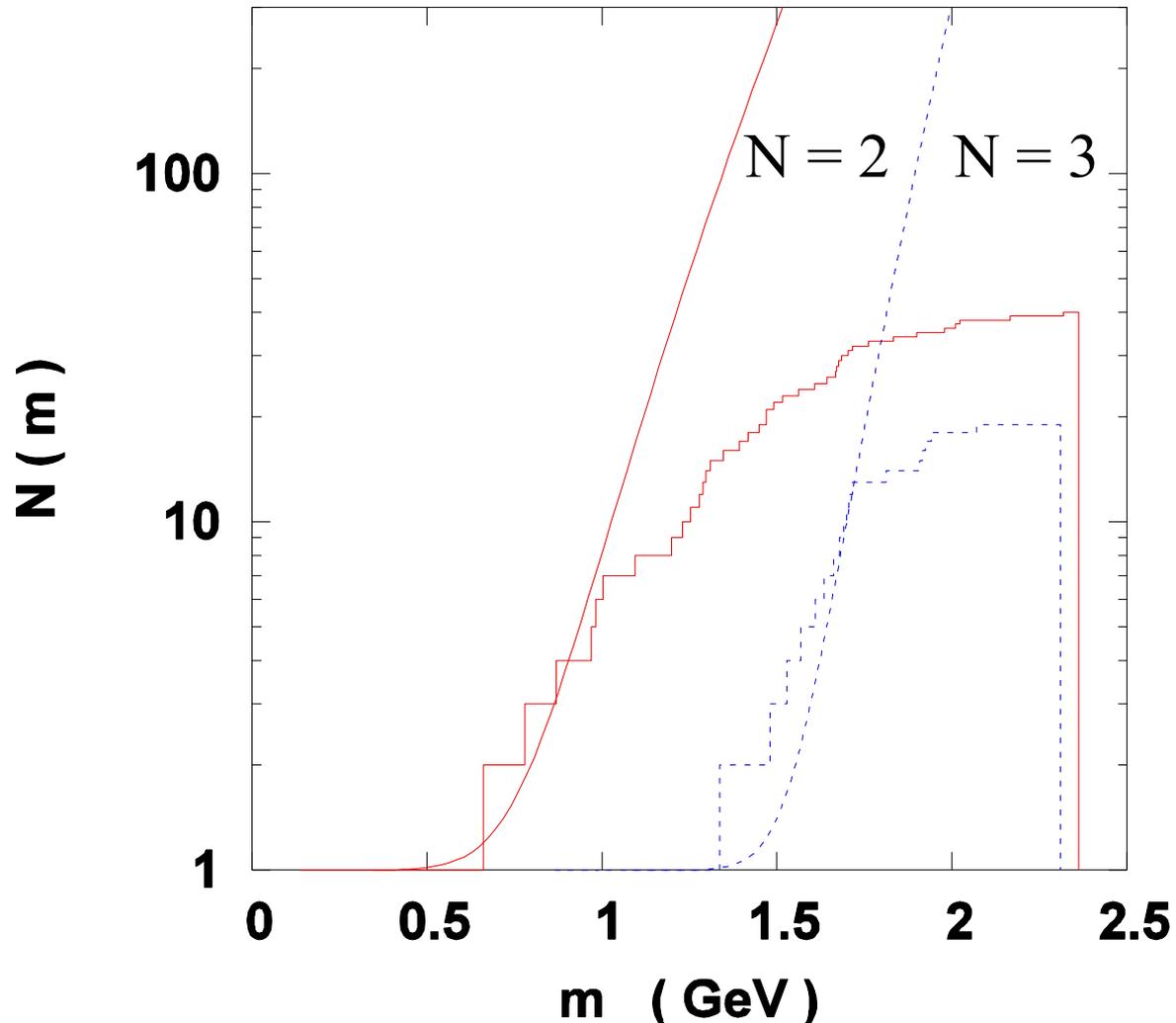
For $N \geq 2$: Tsallis partons \rightarrow Hagedorn hadrons

TEMPERATURE VS. ENERGY



HADRON MASS SPECTRUM FROM X(E)-

FOLDING OF TSALLIS



The hadronization line in stringy quark matter

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J. Cleymans, University of Cape Town, South Africa

- How can be $E / N = 6 T$?
- Stringy corrections to QGP
- High-T equation of state
- The zero pressure line

theory

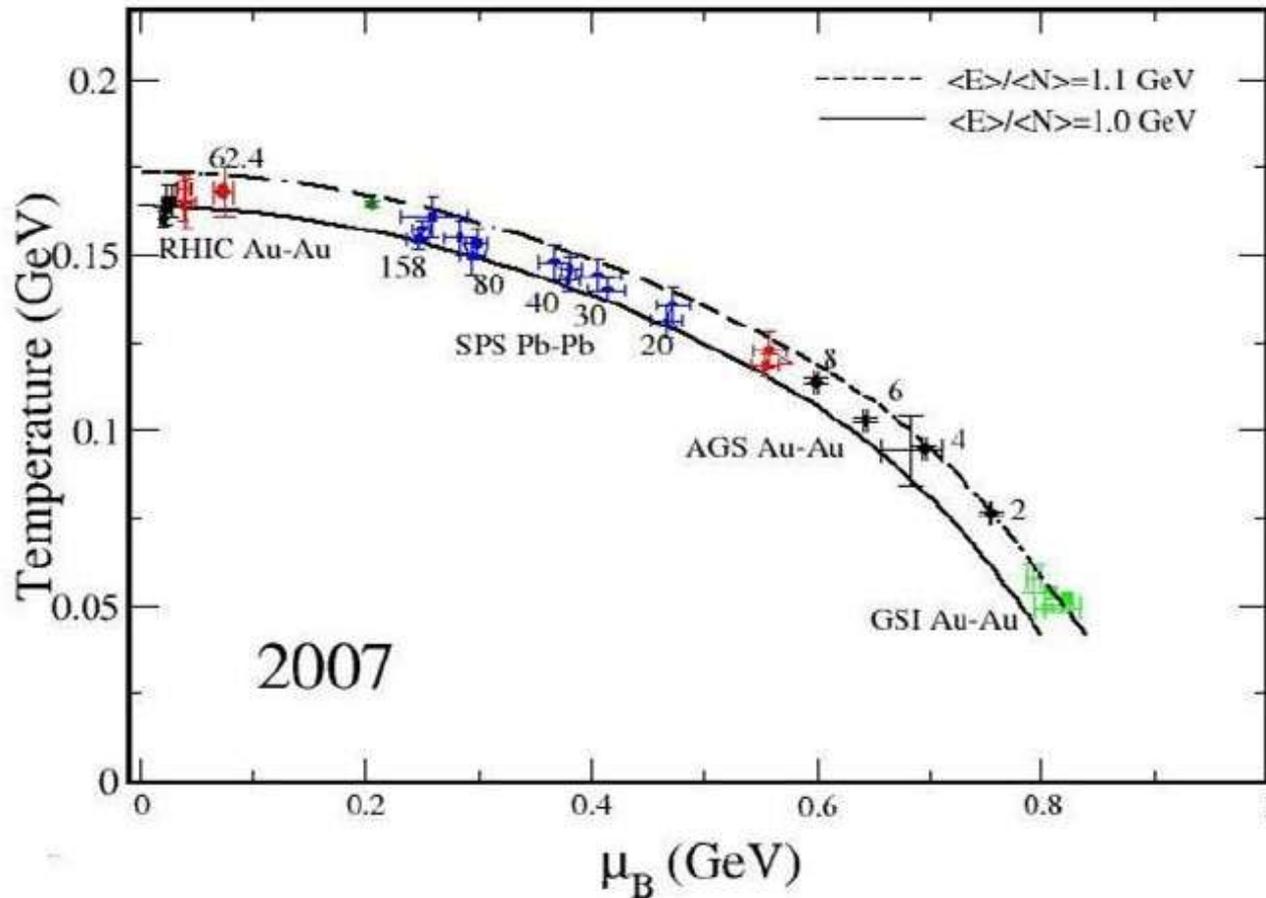
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Statistical Model

Statistical model

Talk: Cleymans



How can be $E / N = 6 T$?

$$T = 167 \text{ MeV}$$

$$E / N = 1 \text{ GeV}$$



Statistical Model: hadronization point around $\mu = 0$ (RHIC, LHC)

How can be $E / N = 6 T$?

$$E / N = m + \frac{3}{2} T$$

$$m = 750 \text{ MeV}$$

Massive hadrons (rho?)

How can be $E / N = 6 T$?

$$E / N = 3T$$

$$m = 0$$

Ideal gas of radiation

How can be $E / N = 6 T$?

$$e = \sigma T^4 + B$$

$$p = \frac{1}{3} \sigma T^4 - B \geq 0$$

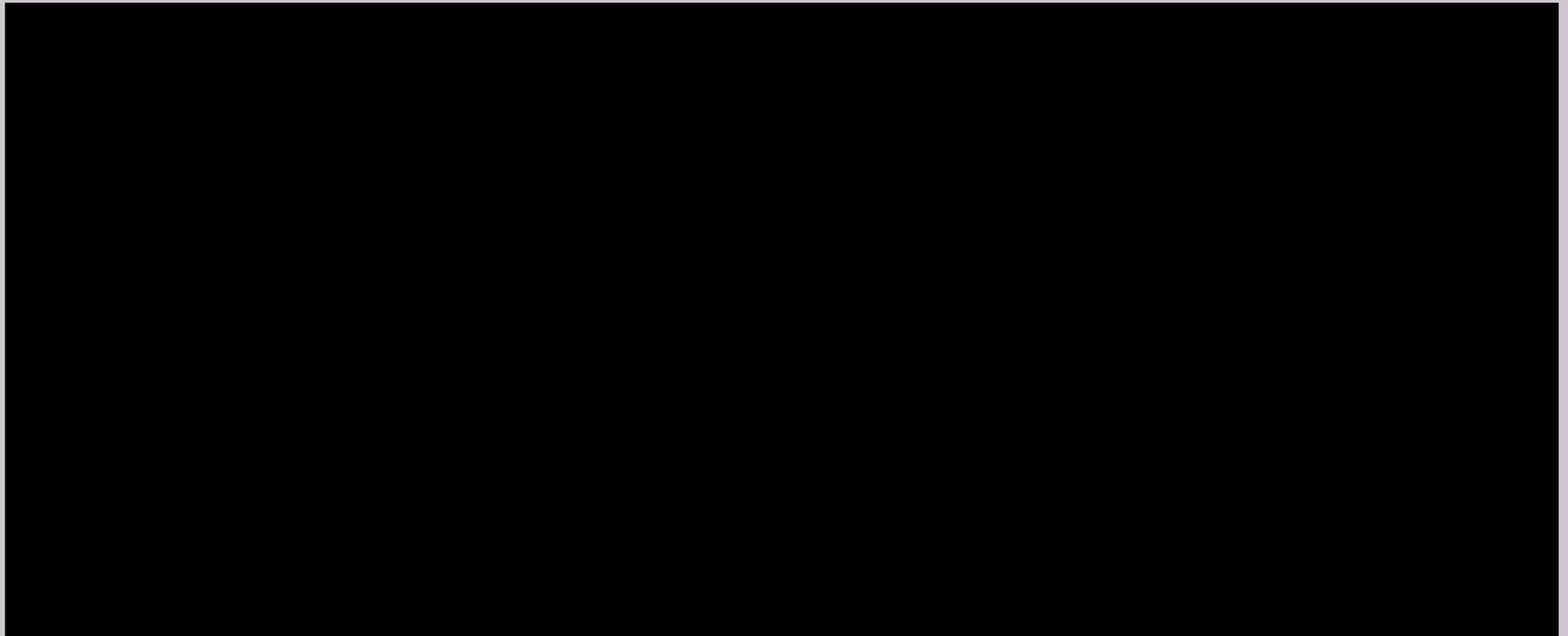
$$n \approx \frac{1}{3} \sigma T^3$$

$$E / N = e / n \leq 4T$$

Bag Model for QGP

How can be $E / N = 6 T$?

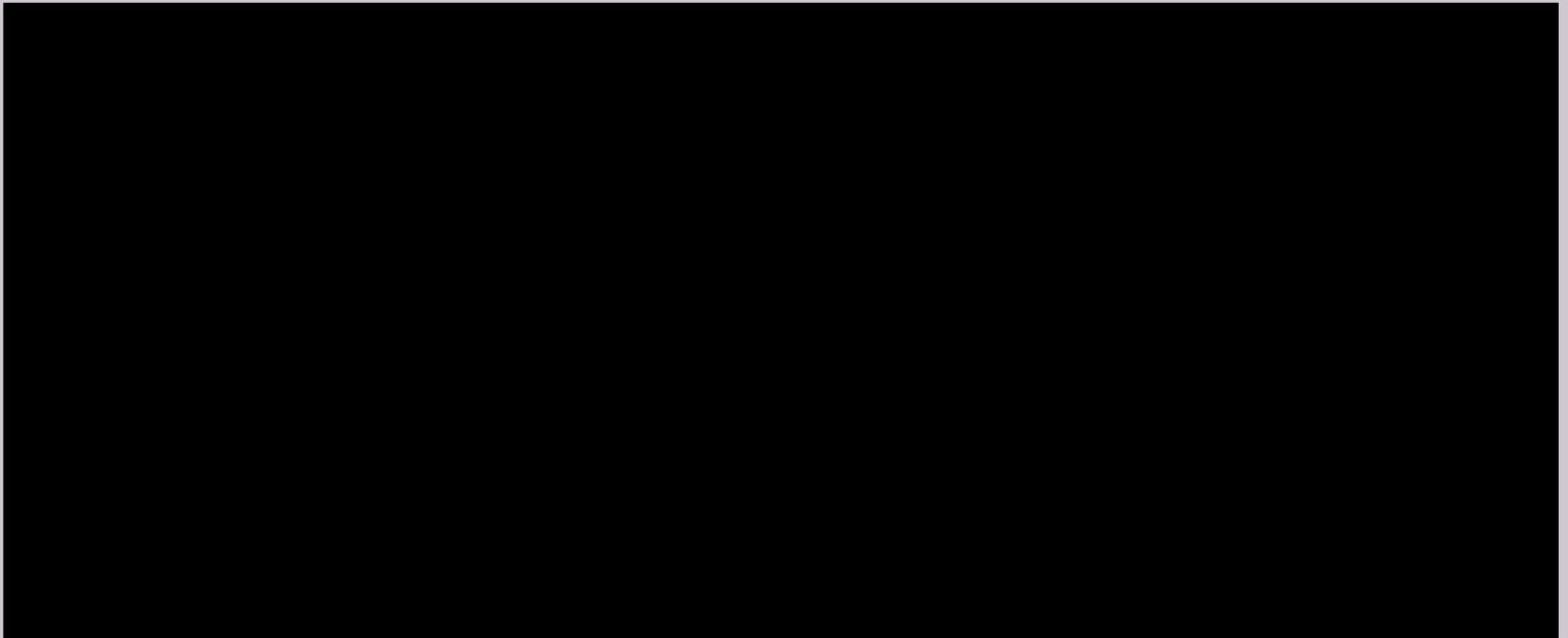
Stringy Massless QGP



How can be $E / N = 6 T$?

Stringy Massless QGP

$$n \cdot \sigma \langle \ell \rangle = \sigma g n^{-1/3} n = \frac{3}{2} A n^{2/3}$$



How can be $E / N = 6 T$?

This is more generally true!

- **correction depends on**

$$c = \sum_i c_i n_i$$

- **constituents are massless**

$$p_{id} = T^4 \cdot \phi\left(\frac{\mu_B}{T}\right)$$

- **Boltzmann approximation is acceptable**

$$p_{id} = T \cdot \sum_i n_i$$

Stringy corrections to QGP

$$c = \sum_i c_i n_i$$

color density

$$f = f_{id} + \frac{1}{1-\gamma} A c^{1-\gamma}$$

free energy density
fractional power

$$\mu_i = \mu_{id} + A c^{-\gamma} c_i = q_i \mu_B$$

chemical potential

$$p = p_{id} - \frac{\gamma}{1-\gamma} A c^{1-\gamma}$$

pressure

$$e = e_{id} + \frac{1}{1-\gamma} A c^{1-\gamma}$$

energy density

Stringy corrections to QGP

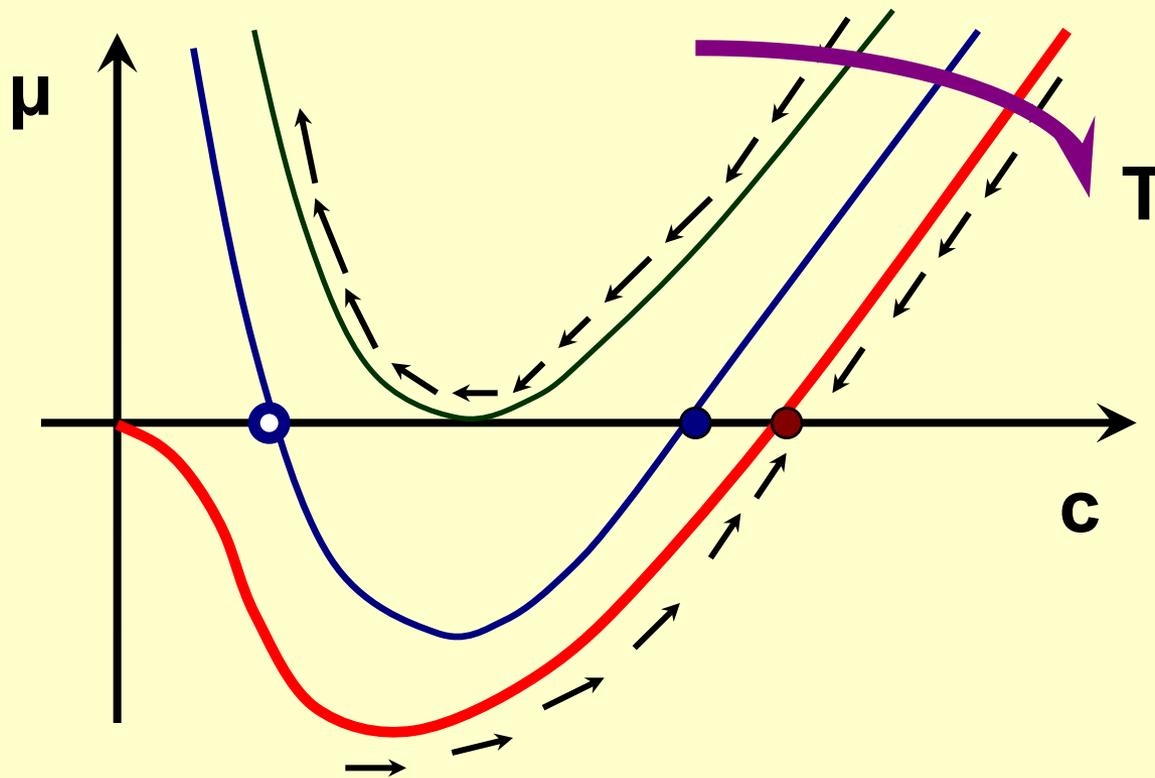
Consistent solution for color density in equilibrium:

$$c = \sum_i c_i d_i n_{id} \left(T, q_i \mu_B - A c_i c^{-\gamma} \right)$$
$$\Rightarrow c = c(T, \mu_B)$$

It exists only beyond an endline in $T - \mu$

Stringy corrections to QGP

Consistent solution for color density in equilibrium:



It exists only beyond an endline in $T - \mu$

Stringy corrections to QGP

One massless component, Boltzmann approximation:

$$c = \frac{d}{\pi^2} T^3 e^{q\mu_B/T} e^{-Ac^{-\gamma}/T}$$

$$z = \frac{A}{T} c^{-\gamma}, \quad \alpha = \frac{A}{T} \left(\frac{d}{\pi^2} T^3 e^{q\mu_B/T} \right)^{-\gamma}$$

$$z = \alpha e^{\gamma z}, \quad z = -\frac{1}{\gamma} W(-\gamma \alpha)$$

The solution is related to Lambert's function

Stringy corrections to QGP

Endline: last possible solution

Boltzmann approximation

$$z = \alpha e^{\gamma z}, \quad 1 = \gamma \alpha e^{\gamma z}$$

$$\Rightarrow z_E = \frac{1}{\gamma} \quad \Rightarrow \quad \alpha_E = \frac{1}{\gamma} e^{-1}$$

$$\mu_B = \frac{3\gamma + 1}{\gamma q} T \ln \frac{T_E}{T}$$

It is near to the zero pressure line for low chemical potential

Stringy corrections to QGP

End temperature: last possible solution at $\mu = 0$

Boltzmann approximation

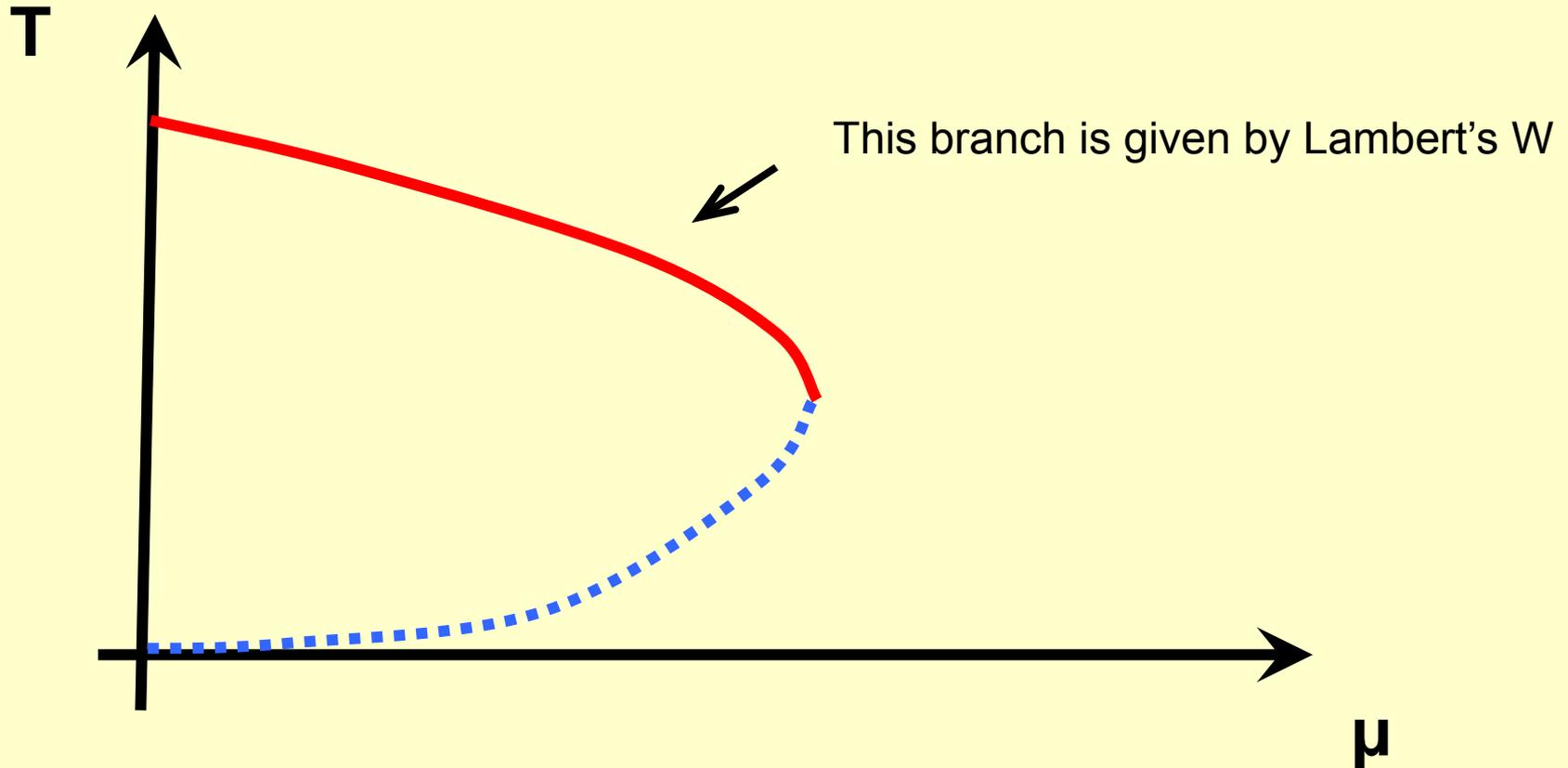
$$T_E = \left[\left(\frac{d}{\pi^2} \right)^{-\gamma} \gamma e A \right]^{\frac{1}{1+3\gamma}}$$

It is near to the zero pressure line for low chemical potential

Stringy corrections to QGP

Endline: last possible solution

Boltzmann approximation



The zero pressure line

Zero pressure in the Boltzmann approximation

$$p = cT \left(1 - \frac{\gamma}{1-\gamma} z \right)$$

$$\Rightarrow z_0 = \frac{1-\gamma}{\gamma} \quad \Rightarrow \quad \alpha_0 = \left(\frac{1}{\gamma} - 1 \right) e^{\gamma-1}$$

$$\mu_B = \frac{3\gamma + 1}{\gamma q} T \ln \frac{T_0}{T}, \quad \frac{T_0}{T_E} = (1 - \gamma)^{-\frac{1}{3\gamma+1}} e^{-\frac{\gamma}{3\gamma+1}}$$

Endline and zero pressure line are relatively close! ($T_0 \approx 1.04 T_E$)

E/N at the zero pressure line

General relations

$$\left. \frac{e}{n} \right|_{p=0} = \frac{\sum [e_{id} + \frac{1}{\gamma} p_{id}]}{\sum n_{id}}$$

$$\left. \frac{e}{n} \right|_{p=m=0} = \left(3 + \frac{1}{\gamma}\right) T \frac{\sum d_i \Phi_i(\tilde{\mu}_i / T)}{\sum d_i \varphi_i(\tilde{\mu}_i / T)}$$

$$\left. \frac{e}{n} \right|_{p=m=0, \text{ Boltzmann}} = \left(3 + \frac{1}{\gamma}\right) T$$

This is independent of the value of the string tension!

High – T equation of state

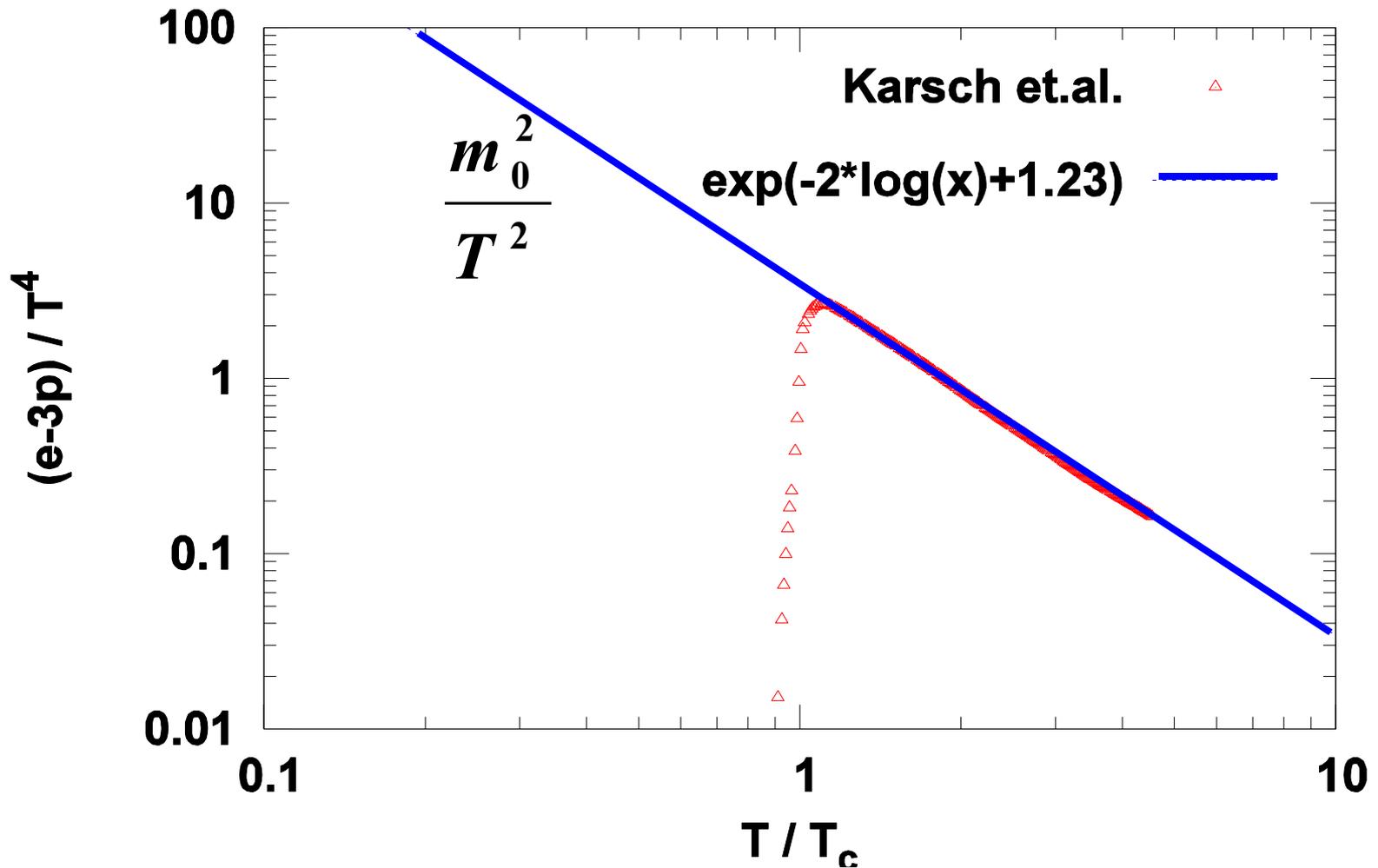
$$\frac{e - 3p}{T^4} = \frac{1 + 3\gamma}{1 - \gamma} \frac{A}{T^4} c^{1-\gamma} \sim T^{-3\gamma-1}$$

lattice QCD:

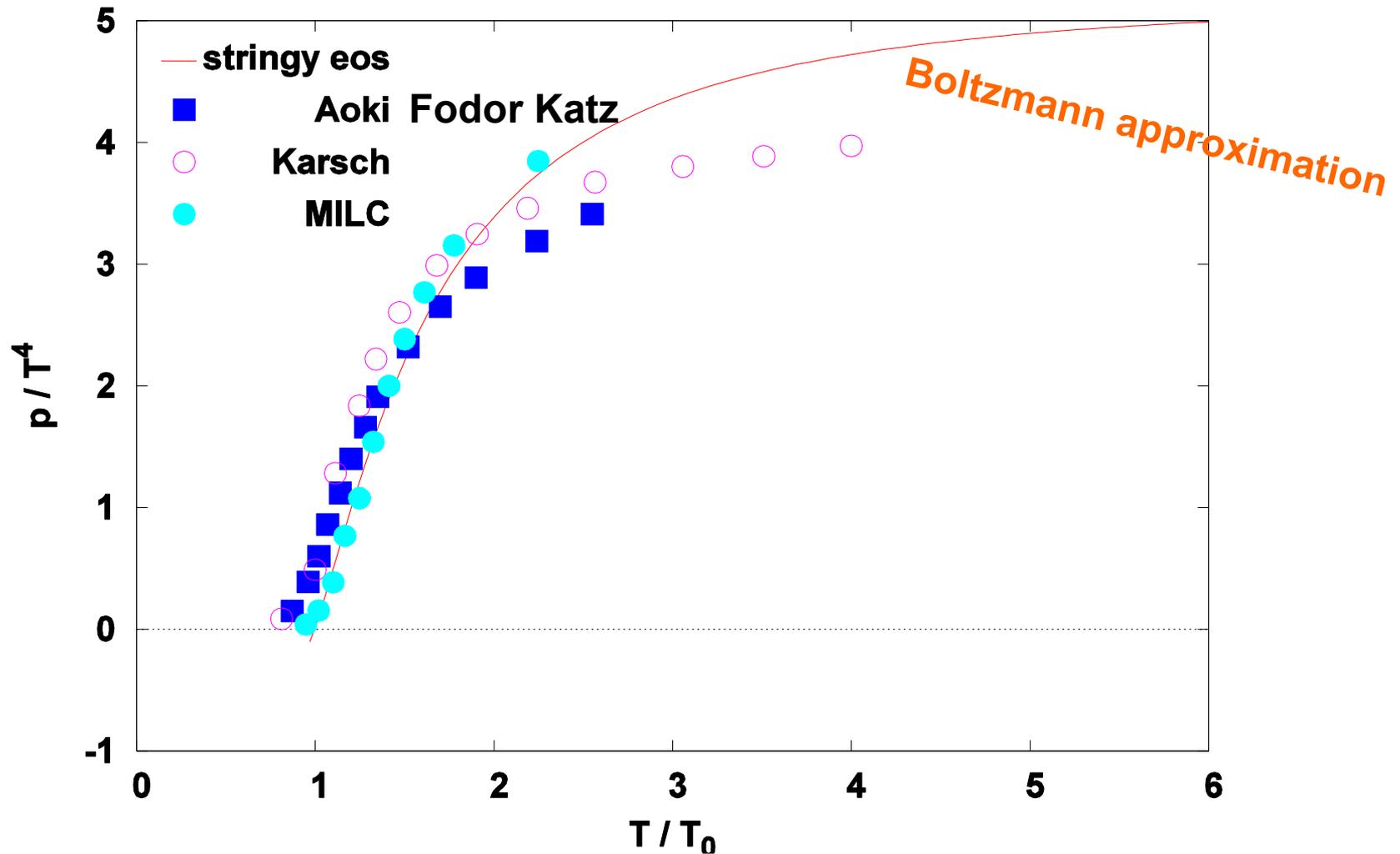
$$3\gamma + 1 = 2 \quad \rightarrow \quad \gamma = \frac{1}{3}$$

High-T behavior of lattice eos

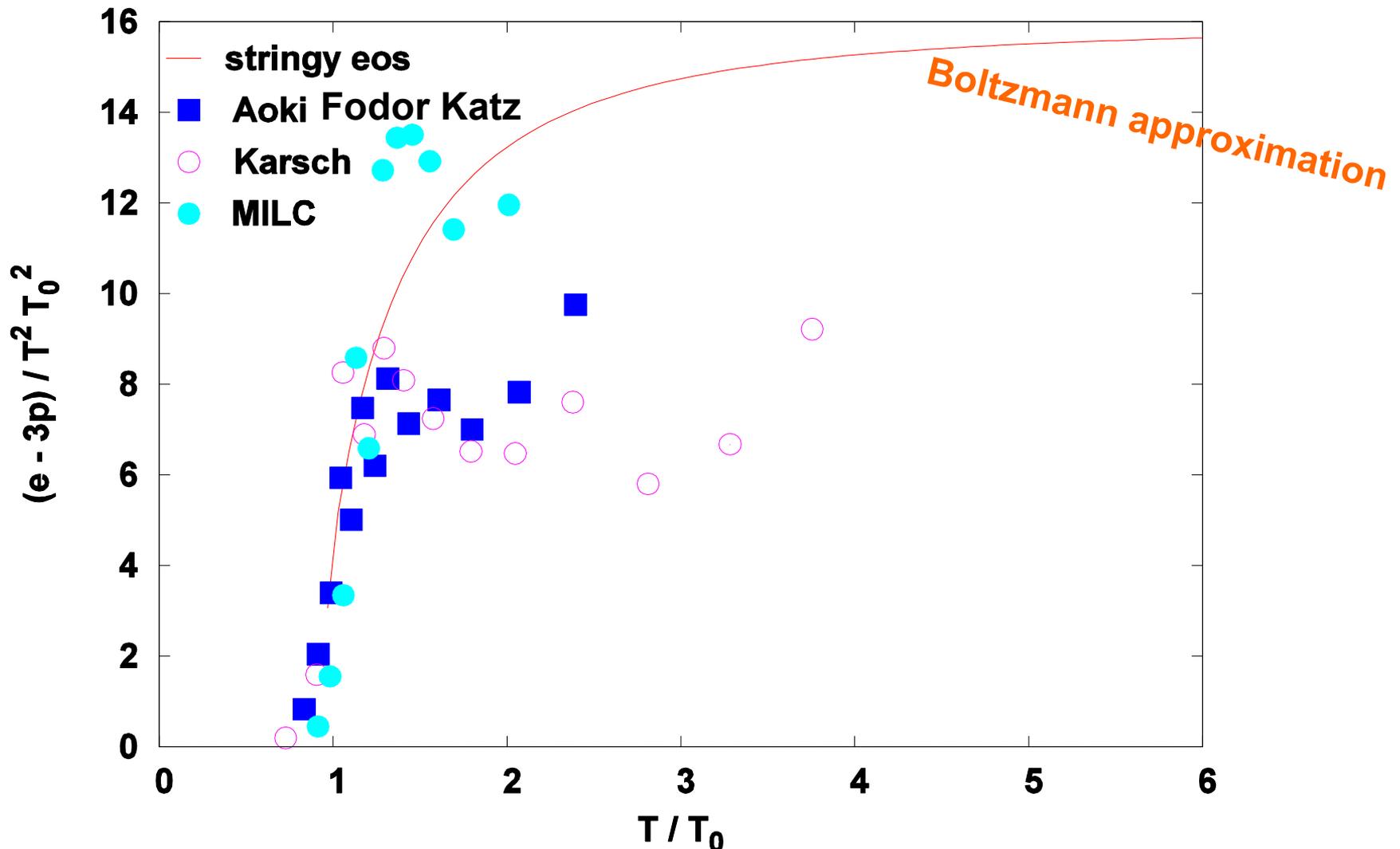
Interaction measure log-log plot



High-T behavior of lattice eos



High-T behavior of lattice eos



Where to stop with the Boltzmann approximation?

Fermi integral \approx Boltzmann integral for negative or zero μ .

$$q\mu_B - Ac^{-\gamma} \leq 0$$

$$\frac{q\mu_B}{T} = \left(3 + \frac{1}{\gamma}\right) \ln \frac{T_E}{T} \leq z_E = \frac{1}{\gamma}$$

$$T_E > T > T_E e^{-1/2} \approx 0.6 T_E$$

The endline does not turn back!

$$T_{\max} = T_E e^{-1}$$

The zero pressure line

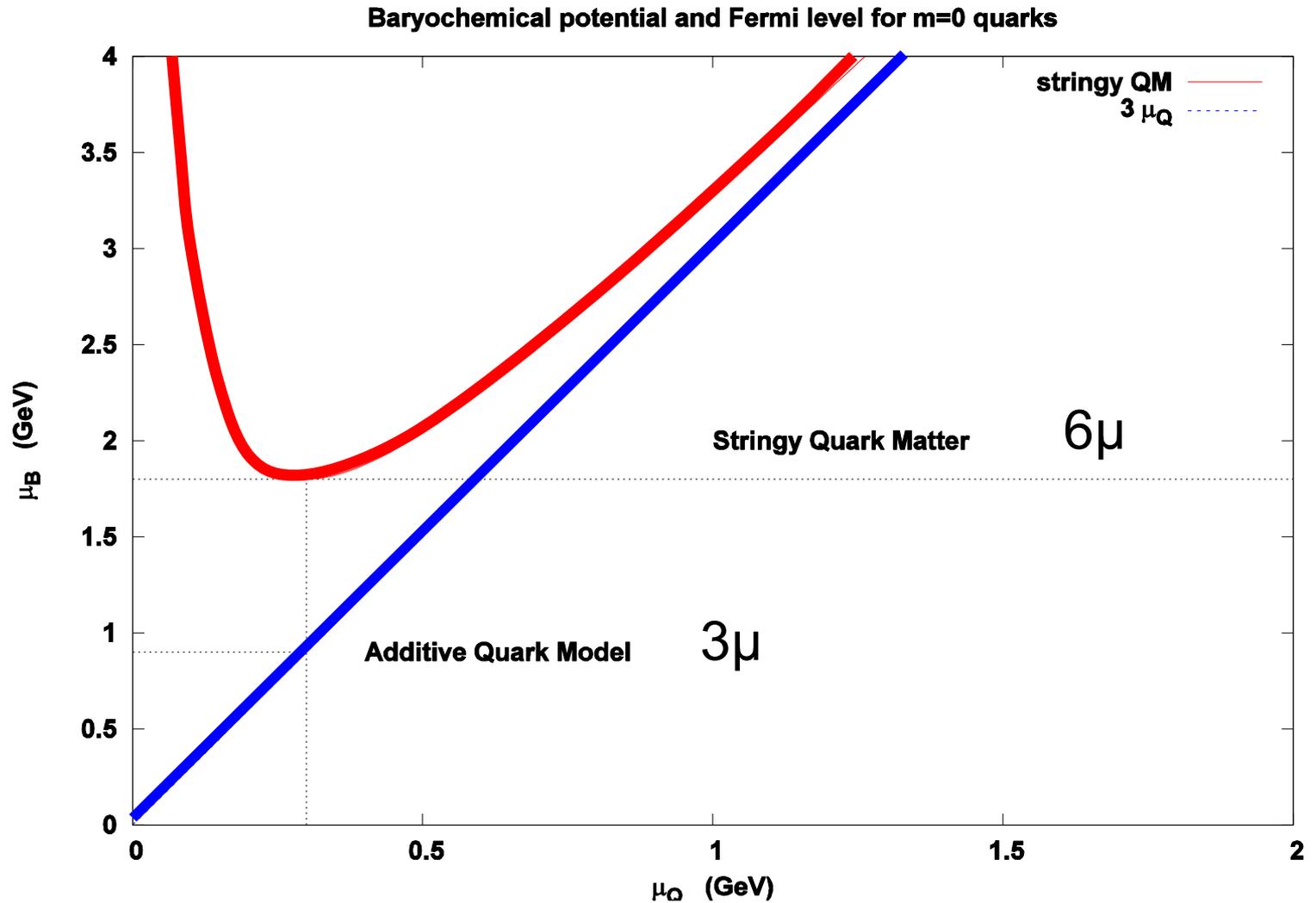
Fermions (quarks only) at T = 0

$$c = \frac{dc_q}{6\pi^2} \left(q\mu_B - Ac^{-1/3}c_q \right)^3$$

$$c = c_q \left[\frac{q\mu_B}{2B} + \sqrt{\frac{q^2\mu_B^2}{4B^2} - \frac{\tilde{A}}{B}} \right]^3, \quad \left. \frac{e}{n} \right|_{p=0, T=0} = q\mu_B$$

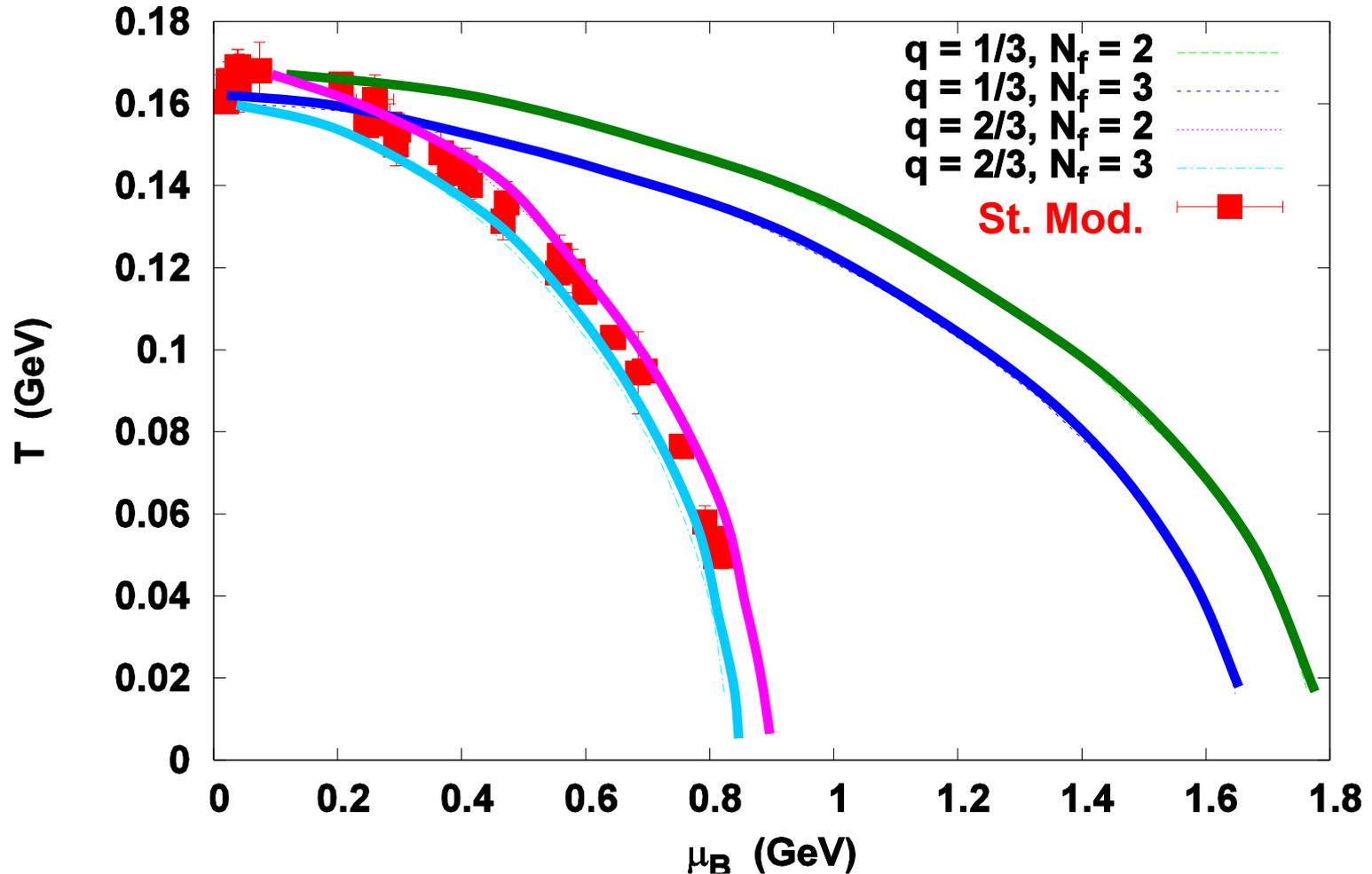
$$\mu_0^Q = \frac{3}{2\sqrt{2}} \mu_E \approx 1.06\mu_E \approx 10.4 T_0^{QGP} \approx 1.74 \text{ GeV}$$

The endpoint at T=0



The zero pressure line

Stringy massless QGP $p = 0$ line ($c_i = 1$, $A = 0.04564 \text{ GeV}^2$)



Discussion





Is acceleration a heat container?